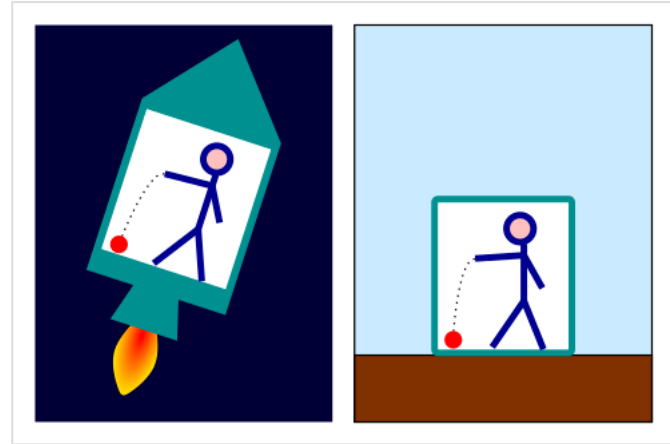




Equivalence principle

The **equivalence principle** is the hypothesis that the observed equivalence of gravitational and inertial mass is a consequence of nature. The weak form, known for centuries, relates to masses of any composition in free fall taking the same trajectories and landing at identical times. The extended form by Albert Einstein requires special relativity to also hold in free fall and requires the weak equivalence to be valid everywhere. This form was a critical input for the development of the theory of general relativity. The strong form requires Einstein's form to work for stellar objects. Highly precise experimental tests of the principle limit possible deviations from equivalence to be very small.



A falling object behaves exactly the same on a planet or in an equivalent accelerating frame of reference.

Concept

In classical mechanics, Newton's equation of motion in a gravitational field, written out in full, is:

$$\text{inertial mass} \times \text{acceleration} = \text{gravitational mass} \times \text{intensity of the gravitational field}$$

Very careful experiments have shown that the inertial mass on the left side and gravitational mass on the right side are numerically equal and independent of the material composing the masses. The equivalence principle is the hypothesis that this numerical equality of inertial and gravitational mass is a consequence of their fundamental identity.^{[1]:32}

The equivalence principle can be considered an extension of the principle of relativity, the principle that the laws of physics are invariant under uniform motion. An observer in a windowless room cannot distinguish between being on the surface of the Earth and being in a spaceship in deep space accelerating at $1g$ and the laws of physics are unable to distinguish these cases.^{[1]:33}

History

Galileo compared different materials experimentally to determine that the acceleration due to gravitation is independent of the amount of mass being accelerated.^[2]

Newton, just 50 years after Galileo, developed the idea that gravitational and inertial mass were different concepts and compared the periods of pendulums composed of different materials to verify that these masses are the same. This form of the equivalence principle became known as "weak equivalence".^[2]

A version of the equivalence principle consistent with special relativity was introduced by Albert Einstein in 1907, when he observed that identical physical laws are observed in two systems, one subject to a constant gravitational field causing acceleration and the other subject to constant acceleration like a rocket far from any gravitational field.^{[3]:152} Since the physical laws are the same, Einstein assumed the gravitational field and the acceleration were "physically equivalent". Einstein stated this hypothesis as:

we ... assume the complete physical equivalence of a gravitational field and a corresponding acceleration of the reference system.

—Einstein, 1907^[4]

In 1911 Einstein demonstrated the power of the equivalence principle by using it to predict that clocks run at different rates in a gravitational potential, and light rays bend in a gravitational field.^{[3]:153} He connected the equivalence principle to his earlier principle of special relativity:

This assumption of exact physical equivalence makes it impossible for us to speak of the absolute acceleration of the system of reference, just as the usual theory of relativity forbids us to talk of the absolute velocity of a system; and it makes the equal falling of all bodies in a gravitational field seem a matter of course.

—Einstein, 1911^[5]

Immediately after completing his work^{[6]:111} on a theory of gravity (known as general relativity) and in later years Einstein recalled the role of the equivalence principle:

The breakthrough came suddenly one day. I was sitting on a chair in my patent office in Bern. Suddenly a thought struck me: If a man falls freely, he would not feel his weight. I was taken aback. This simple thought experiment made a deep impression on me. This led me to the theory of gravity.

—Einstein, 1922^[7]

Since Einstein developed general relativity, there was a need to develop a framework to test the theory against other possible theories of gravity compatible with special relativity. This was developed by Robert Dicke as part of his program to test general relativity. Two new principles were suggested, the so-called Einstein equivalence principle and the strong equivalence principle, each of which assumes the weak equivalence principle as a starting point. These are discussed below.

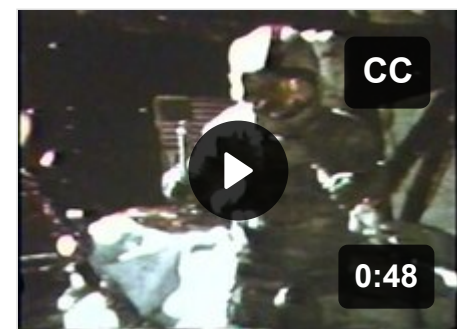
Definitions

Three main forms of the equivalence principle are in current use: weak (Galilean), Einsteinian, and strong.^{[8]:6} Some studies also create finer divisions or slight alternative.^{[9][10]}

Weak equivalence principle

The weak equivalence principle, also known as the universality of free fall or the Galilean equivalence principle can be stated in many ways. The strong equivalence principle, a generalization of the weak equivalence principle, includes astronomic bodies with gravitational self-binding energy.^[11] Instead, the weak equivalence principle assumes falling bodies are self-bound by non-gravitational forces only (e.g. a stone). Either way:

- "All uncharged, freely falling test particles follow the same trajectories, once an initial position and velocity have been prescribed".^{[8]:6}
- "... in a uniform gravitational field all objects, regardless of their composition, fall with precisely the same acceleration." "The weak equivalence principle implicitly assumes that the falling objects are bound by non-gravitational forces."^[11]



During the Apollo 15 mission in 1971, astronaut David Scott showed that Galileo was right: acceleration is the same for all bodies subject to gravity on the Moon, even for a hammer and a feather.

- "... in a gravitational field the acceleration of a test particle is independent of its properties, including its rest mass."^[12]
- Mass (measured with a balance) and weight (measured with a scale) are locally in identical ratio for all bodies (the opening page to Newton's *Philosophiæ Naturalis Principia Mathematica*, 1687).

Uniformity of the gravitational field eliminates measurable tidal forces originating from a radial divergent gravitational field (e.g., the Earth) upon finite sized physical bodies.

Einstein equivalence principle

What is now called the "Einstein equivalence principle" states that the weak equivalence principle holds, and that:

the outcome of any local, non-gravitational test experiment is independent of the experimental apparatus' velocity relative to the gravitational field and is independent of where and when in the gravitational field the experiment is performed.^[13]

Here *local* means that experimental setup must be small compared to variations in the gravitational field, called tidal forces. The *test* experiment must be small enough so that its gravitational potential does not alter the result.

The two additional constraints added to the weak principle to get the Einstein form – (1) the independence of the outcome on relative velocity (local Lorentz invariance) and (2) independence of "where" known as (local positional invariance) – have far reaching consequences. With these constraints alone Einstein was able to predict the gravitational redshift.^[13] Theories of gravity that obey the Einstein equivalence principle must be "metric theories", meaning that trajectories of freely falling bodies are geodesics of symmetric metric.^{[14]:9}

Around 1960 Leonard I. Schiff conjectured that any complete and consistent theory of gravity that embodies the weak equivalence principle implies the Einstein equivalence principle; the conjecture can't be proven but has several plausibility arguments in its favor.^{[14]:20} Nonetheless, the two principles are tested with very different kinds of experiments.

The Einstein equivalence principle has been criticized as imprecise, because there is no universally accepted way to distinguish gravitational from non-gravitational experiments (see for instance Hadley^[15] and Durand^[16]).

Strong equivalence principle

The strong equivalence principle applies the same constraints as the Einstein equivalence principle, but allows the freely falling bodies to be massive gravitating objects as well as test particles.^[8] Thus this is a version of the equivalence principle that applies to objects that exert a gravitational force on themselves, such as stars, planets, black holes or Cavendish experiments. It requires that the gravitational constant be the same everywhere in the universe^{[14]:49} and is incompatible with a fifth force. It is much more restrictive than the Einstein equivalence principle.

Like the Einstein equivalence principle, the strong equivalence principle requires gravity is geometrical by nature, but in addition it forbids any extra fields, so the metric alone determines all of the effects of gravity. If an observer measures a patch of space to be flat, then the strong equivalence principle suggests that it is absolutely equivalent to any other patch of flat space elsewhere in the universe. Einstein's theory of general

relativity (including the cosmological constant) is thought to be the only theory of gravity that satisfies the strong equivalence principle. A number of alternative theories, such as Brans–Dicke theory and the Einstein-aether theory add additional fields.^[8]

Active, passive, and inertial masses

Some of the tests of the equivalence principle use names for the different ways mass appears in physical formulae. In nonrelativistic physics three kinds of mass can be distinguished:^[14]

1. Inertial mass intrinsic to an object, the sum of all of its mass–energy.
2. Passive mass, the response to gravity, the object's weight.
3. Active mass, the mass that determines the objects gravitational effect.

By definition of active and passive gravitational mass, the force on \mathbf{M}_1 due to the gravitational field of \mathbf{M}_0 is:

$$\mathbf{F}_1 = \frac{M_0^{\text{act}} M_1^{\text{pass}}}{r^2}$$

Likewise the force on a second object of arbitrary mass₂ due to the gravitational field of mass₀ is:

$$\mathbf{F}_2 = \frac{M_0^{\text{act}} M_2^{\text{pass}}}{r^2}$$

By definition of inertial mass:

$$\mathbf{F} = m^{\text{inert}} \mathbf{a}$$

if m_1 and m_2 are the same distance r from m_0 then, by the weak equivalence principle, they fall at the same rate (i.e. their accelerations are the same).

$$a_1 = \frac{F_1}{m_1^{\text{inert}}} = a_2 = \frac{F_2}{m_2^{\text{inert}}}$$

Hence:

$$\frac{M_0^{\text{act}} M_1^{\text{pass}}}{r^2 m_1^{\text{inert}}} = \frac{M_0^{\text{act}} M_2^{\text{pass}}}{r^2 m_2^{\text{inert}}}$$

Therefore:

$$\frac{M_1^{\text{pass}}}{m_1^{\text{inert}}} = \frac{M_2^{\text{pass}}}{m_2^{\text{inert}}}$$

In other words, passive gravitational mass must be proportional to inertial mass for objects, independent of their material composition if the weak equivalence principle is obeyed.

The dimensionless Eötvös-parameter or Eötvös ratio $\eta(\mathbf{A}, \mathbf{B})$ is the difference of the ratios of gravitational and inertial masses divided by their average for the two sets of test masses "A" and "B".

$$\eta(\mathbf{A}, \mathbf{B}) = 2 \frac{\left(\frac{m_{\text{pass}}}{m_{\text{inert}}} \right)_A - \left(\frac{m_{\text{pass}}}{m_{\text{inert}}} \right)_B}{\left(\frac{m_{\text{pass}}}{m_{\text{inert}}} \right)_A + \left(\frac{m_{\text{pass}}}{m_{\text{inert}}} \right)_B}.$$

Values of this parameter are used to compare tests of the equivalence principle.^{[14]:10}

A similar parameter can be used to compare passive and active mass. By Newton's third law of motion:

$$F_1 = \frac{M_0^{\text{act}} M_1^{\text{pass}}}{r^2}$$

must be equal and opposite to

$$F_0 = \frac{M_1^{\text{act}} M_0^{\text{pass}}}{r^2}$$

It follows that:

$$\frac{M_0^{\text{act}}}{M_0^{\text{pass}}} = \frac{M_1^{\text{act}}}{M_1^{\text{pass}}}$$

In words, passive gravitational mass must be proportional to active gravitational mass for all objects. The difference,

$$S_{0,1} = \frac{M_0^{\text{act}}}{M_0^{\text{pass}}} - \frac{M_1^{\text{act}}}{M_1^{\text{pass}}}$$

is used to quantify differences between passive and active mass.^[17]

Experimental tests

Tests of the weak equivalence principle

Tests of the weak equivalence principle are those that verify the equivalence of gravitational mass and inertial mass. An obvious test is dropping different objects and verifying that they land at the same time. Historically this was the first approach, though probably not by Galileo's Leaning Tower of Pisa experiment^{[18]:19–21} but earlier by Simon Stevin^[19] who dropped lead balls of different masses off the Delft churchtower and listened for the sound they made on a wooden plank.

Isaac Newton measured the period of pendulums made with different materials as an alternative test giving the first precision measurements.^[2] Loránd Eötvös's approach in 1908 used a very sensitive torsion balance to give precision approaching 1 in a billion. Modern experiments have improved this by another factor of a million.

A popular exposition of this measurement was done on the Moon by David Scott in 1971. He dropped a falcon feather and a hammer at the same time, showing on video^[20] that they landed at the same time.

Chronology of weak equivalence principles tests^[21]

Year	Investigator	Sensitivity	Method
500?	<u>John Philoponus</u> ^[22]	"small"	Drop tower
1585	<u>Simon Stevin</u> ^{[23][19]}	5×10^{-2}	Drop tower
1590?	<u>Galileo Galilei</u> ^{[24][21]:91}	2×10^{-3}	Pendulum, drop tower
1686	<u>Isaac Newton</u> ^{[25][21]:91}	10^{-3}	Pendulum
1832	<u>Friedrich Wilhelm Bessel</u> ^{[26][21]:91}	2×10^{-5}	Pendulum
1908 (1922)	<u>Loránd Eötvös</u> ^{[27][21]:92}	2×10^{-9}	Torsion balance
1910	<u>Southerns</u> ^{[28][21]:91}	5×10^{-6}	Pendulum
1918	<u>Zeeman</u> ^{[29][21]:91}	3×10^{-8}	Torsion balance
1923	<u>Potter</u> ^{[30][21]:91}	3×10^{-6}	Pendulum
1935	<u>Renner</u> ^{[31][21]:92}	2×10^{-9}	Torsion balance
1964	<u>Roll, Krotkov, Dicke</u> ^[32]	3×10^{-11}	Torsion balance
1972	<u>Braginsky, Panov</u> ^{[33][21]:92}	10^{-12}	Torsion balance
1976	<u>Shapiro, et al.</u> ^{[34][21]:92}	10^{-12}	Lunar laser ranging
1979	<u>Keiser, Faller</u> ^{[35][21]:93}	4×10^{-11}	Fluid support
1987	<u>Niebauer, et al.</u> ^{[36][21]:95}	10^{-10}	Drop tower
1989	<u>Stubbs, et al.</u> ^{[37][21]:93}	10^{-11}	Torsion balance
1990	<u>Adelberger, Eric G.; et al.</u> ^{[38][21]:95}	10^{-12}	Torsion balance
1999	<u>Baessler, et al.</u> ^{[39][40]}	5×10^{-14}	Torsion balance
2008	<u>Schlamming, et al.</u> ^[41]	10^{-13}	Torsion balance
2017	<u>MICROSCOPE</u> ^{[42][43]}	10^{-15}	Earth orbit

Experiments are still being performed at the University of Washington which have placed limits on the differential acceleration of objects towards the Earth, the Sun and towards dark matter in the Galactic Center.^[44] Future satellite experiments^[45] – Satellite Test of the Equivalence Principle^[46] and Galileo Galilei – will test the weak equivalence principle in space, to much higher accuracy.^[47]

With the first successful production of antimatter, in particular anti-hydrogen, a new approach to test the weak equivalence principle has been proposed. Experiments to compare the gravitational behavior of matter and antimatter are currently being developed.^[48]

Proposals that may lead to a quantum theory of gravity such as string theory and loop quantum gravity predict violations of the weak equivalence principle because they contain many light scalar fields with long Compton wavelengths, which should generate fifth forces and variation of the fundamental constants. Heuristic arguments suggest that the magnitude of these equivalence principle violations could be in the 10^{-13} to 10^{-18} range.^[49]

Currently envisioned tests of the weak equivalence principle are approaching a degree of sensitivity such that *non-discovery* of a violation would be just as profound a result as discovery of a violation. Non-discovery of equivalence principle violation in this range would suggest that gravity is so fundamentally different from other forces as to require a major reevaluation of current attempts to unify gravity with the other forces of nature. A positive detection, on the other hand, would provide a major guidepost towards unification.^[49]

Tests of the Einstein equivalence principle

In addition to the tests of the weak equivalence principle, the Einstein equivalence principle requires testing the local Lorentz invariance and local positional invariance conditions.

Testing local Lorentz invariance amounts to testing special relativity, a theory with vast number of existing tests.^{[14]:12} Nevertheless, attempts to look for quantum gravity require even more precise tests. The modern tests include looking for directional variations in the speed of light (called "clock anisotropy tests") and new forms of the Michelson-Morley experiment. The anisotropy measures less than one part in 10^{-20} .^{[14]:14}

Testing local positional invariance divides in to tests in space and in time.^{[14]:17} Space-based tests use measurements of the gravitational redshift, the classic is the Pound–Rebka experiment in the 1960s. The most precise measurement was done in 1976 by flying a hydrogen maser and comparing it to one on the ground. The Global positioning system requires compensation for this redshift to give accurate position values.

Time-based tests search for variation of dimensionless constants and mass ratios.^[50] For example, Webb et al.^[51] reported detection of variation (at the 10^{-5} level) of the fine-structure constant from measurements of distant quasars. Other researchers dispute these findings.^[52]

The present best limits on the variation of the fundamental constants have mainly been set by studying the naturally occurring Oklo natural nuclear fission reactor, where nuclear reactions similar to ones we observe today have been shown to have occurred underground approximately two billion years ago. These reactions are extremely sensitive to the values of the fundamental constants.

Tests of changes in fundamental constants^{[14]:19}

Constant	Year	Method	Limit on fractional change per year
<u>weak interaction constant</u>	1976	Oklo	10^{-11}
<u>fine-structure constant</u>	1976	Oklo	10^{-16}
<u>electron–proton mass ratio</u>	2002	quasars	10^{-15}

Tests of the strong equivalence principle

The strong equivalence principle can be tested by 1) finding orbital variations in massive bodies (Sun-Earth-Moon), 2) variations in the gravitational constant (*G*) depending on nearby sources of gravity or on motion, or 3) searching for a variation of Newton's gravitational constant over the life of the universe^{[14]:47}

Orbital variations due to gravitational self-energy should cause a "polarization" of solar system orbits called the Nordtvedt effect. This effect has been sensitively tested by the Lunar Laser Ranging Experiment.^{[53][54]} Up to the limit of one part in 10^{13} there is no Nordtvedt effect.

A tight bound on the effect of nearby gravitational fields on the strong equivalence principle comes from modeling the orbits of binary stars and comparing the results to pulsar timing data.^{[14]:49} In 2014, astronomers discovered a stellar triple system containing a millisecond pulsar PSR J0337+1715 and two white

dwarfs orbiting it. The system provided them a chance to test the strong equivalence principle in a strong gravitational field with high accuracy.^{[55][56][57][58]}

Most alternative theories of gravity predict a change in the gravity constant over time. Studies of Big Bang nucleosynthesis, analysis of pulsars, and the lunar laser ranging data have shown that *G* cannot have varied by more than 10% since the creation of the universe. The best data comes from studies of the ephemeris of Mars, based on three successive NASA missions, Mars Global Surveyor, Mars Odyssey, and Mars Reconnaissance Orbiter.^{[14]:50}

See also

- Classical mechanics
- Eötvös experiment
- Einstein's thought experiments
- Gauge gravitation theory
- General covariance
- Mach's principle
- Tests of general relativity
- Unsolved problems in astronomy
- Unsolved problems in physics

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Further reading

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- Misner, Charles W.; Thorne, Kip S.; and Wheeler, John A.; *Gravitation*, New York: W. H. Freeman and Company, 1973, Chapter 16 discusses the equivalence principle.
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- Friedman, Michael; *Foundations of Space-Time Theories*, Princeton, New Jersey: Princeton University Press, 1983. Chapter V discusses the equivalence principle.

External links

- Gravity and the principle of equivalence – The Feynman Lectures on Physics (https://feynmanlectures.caltech.edu/II_42.html#Ch42-S5)
 - Introducing The Einstein Principle of Equivalence (<https://web.archive.org/web/20080515172036/http://www.phy.syr.edu/courses/modules/LIGHTCONE/equivalence.html>) from Syracuse University
 - The Equivalence Principle (<http://www.mathpages.com/rr/s5-06/5-06.htm>) at MathPages
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 - "...Physicists in Germany have used an atomic interferometer to perform the most accurate ever test of the equivalence principle at the level of atoms..." (<https://physicsworld.com/a/equivalence-principle-passe-s-atomic-test/>)
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